

BEYOND POPULAR SCIENCE



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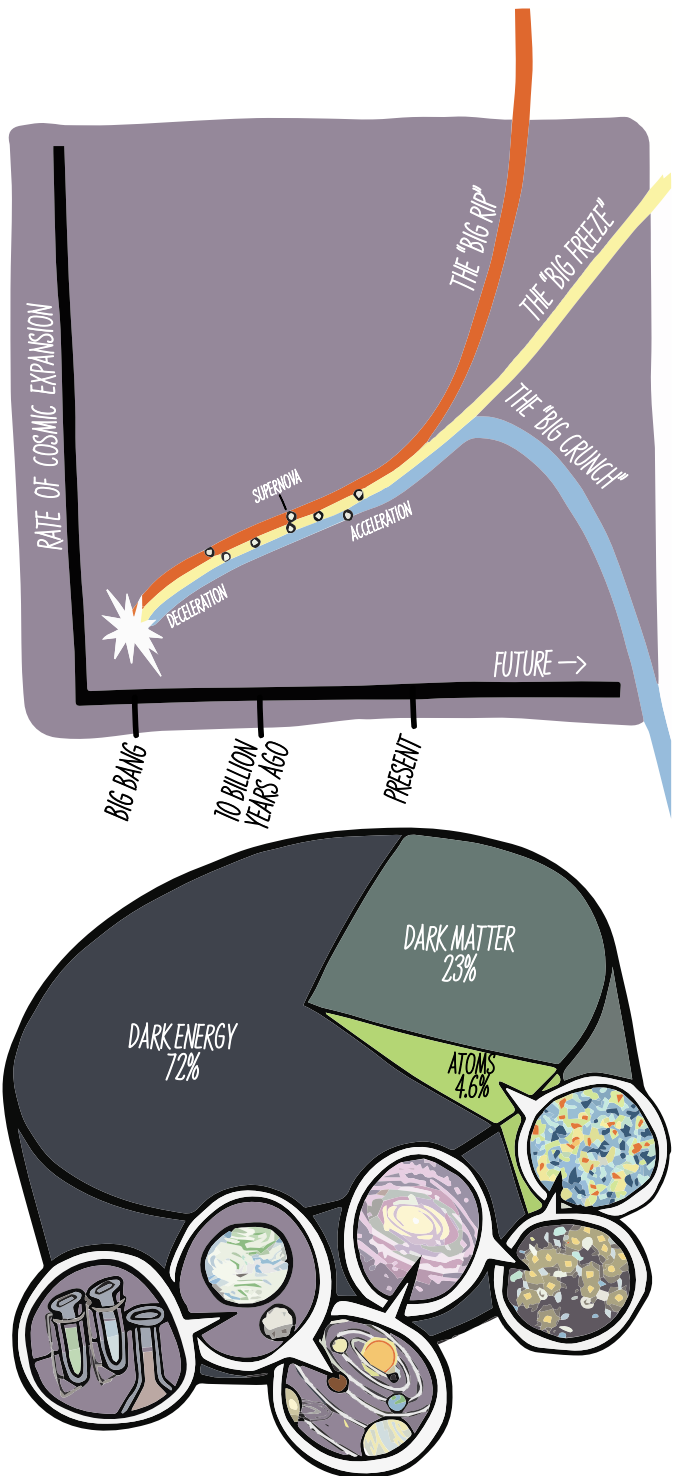
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**Dark Energies
Are Pushing
Us Apart**

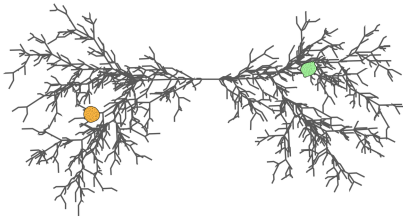
Top (Cosmic Expansion Scenarios): The universe's rate of expansion is shown across time, beginning with an early deceleration due to gravity, followed by the observed acceleration linked to dark energy. Possible future outcomes are indicated: the *Big Crunch* (re-collapse), the *Big Freeze* (endless cooling expansion), and the *Big Rip* (runaway acceleration tearing structures apart). Supernova data points provide the evidence anchoring this curve.

Bottom (Cosmic Energy Composition): The universe's mass-energy content is divided into dark energy (72%), dark matter (23%), and atoms (4.6%), with small additional contributions from neutrinos and photons. Within this narrow slice of atomic matter we can zoom in further: the large-scale structures of the cosmos → stars and galaxies → planetary systems → laboratory beakers. This exponentially dropping sequence highlights how minuscule the portion we directly observe is compared with the vast domain over which the same physical laws are successfully extrapolated.



Dark Energies Are Pushing Us Apart

Observations of distant supernovae in the 1990s revealed that the universe's expansion is accelerating, contradicting earlier models that predicted gravity would slow cosmic expansion. This acceleration requires dark energy, an unknown component comprising about 70% of the universe's energy density. Dark energy counteracts gravity at cosmological scales, manifesting as either a cosmological constant in Einstein's equations or a dynamical field.



TYPE IA SUPERNOVAE ◦ STANDARD CANDLES ◦ HUBBLE
DIAGRAM ◦ COSMIC ACCELERATION DISCOVERY ◦ DARK
ENERGY ◦ COSMOLOGICAL CONSTANT Λ ◦ DISTANCE
MODULUS ◦ REDSHIFT-LUMINOSITY ◦ Λ CDM MODEL ◦ CMB
CONFIRMATION ◦ 68% DARK ENERGY

“The simplest calculation involves summing the quantum mechanical zero-point energies of all the fields known in Nature. This gives an answer about 120 orders of magnitude higher than the upper limits on Λ set by cosmological observations. This is probably the worst theoretical prediction in the history of physics! Nobody knows how to make sense of this result. Some physics mechanism must exist that makes the cosmological constant very small.”

– Hobson, Efstathiou & Lasenby, 2006

Dark Energies Are Pushing Us Apart

In the late 1990s, two independent collaborations studying distant Type Ia supernovae—the Supernova Cosmology Project led by Saul Perlmutter and the High-Z Supernova Search Team led by Adam Riess and Brian Schmidt—reported that these stellar explosions were fainter than anticipated. Their apparent brightness suggested they were farther away than a purely decelerating cosmological model would indicate. This unexpected outcome signalled an accelerating rate of cosmic expansion.

Early theoretical groundwork on cosmic expansion stemmed from solutions to Einstein's field equations by Alexander Friedmann in the 1920s and Georges Lemaître in the 1930s. Although Einstein initially introduced a cosmological constant Λ to force a static universe, Hubble's observations of galactic recession (1929) established expansion as a fact. Decades later, detailed measurements of supernova luminosities confirmed that this expansion was not slowing down, but accelerating.

Recognition of this discovery came in 2011, when the Nobel Prize in Physics was awarded to Perlmutter, Riess, and Schmidt for uncovering cosmic acceleration. Additional evidence arose from measurements of cosmic microwave background anisotropies and large-scale galaxy distributions, all converging on the conclusion that the expansion of the universe is not merely continuing but speeding up.

Stars are thermonuclear engines balanced between gravitational contraction and fusion pressure. This pressure is sustained by nuclear fusion: the conversion of light elements into heavier ones, accompanied by energy release. Low-mass stars fuse hydrogen through the proton-proton chain; massive stars use the CNO (Carbon-Nitrogen-Oxygen) cycle (Bethe, 1939) at their higher core temperatures.

Stellar evolution proceeds as nuclear fuel depletes. Once hydrogen fusion subsides, gravity compresses the core, raising the temperature and enabling new fusion stages. Helium is converted into carbon and oxygen, followed by carbon, neon, oxygen, and silicon burning in increasingly rapid succession. These stages produce successively heavier elements up to iron. Fusion beyond iron is endothermic, and no further energy can be extracted from nuclear reactions.

Stellar fate depends on initial mass. Stars below ~ 8 solar masses cannot fuse heavy elements. Their cores contract into white dwarfs composed of carbon and oxygen, while their outer layers are ejected as planetary nebulae. White dwarfs are stabilised by electron degeneracy pressure up to the Chandrasekhar limit (Chandrasekhar, 1931), $M_{\text{Ch}} \approx 1.4M_{\odot}$ (the symbol \odot denotes the Sun).

Heavier stars burn through all stages to iron cores. The core collapses in milliseconds when fusion ceases. Infalling material rebounds off the dense core and triggers an explosion that ejects the outer layers. This core-collapse supernova leaves behind a neutron star or, if massive enough, a black hole.

Supernovae are classified according to their spectral features. Type II supernovae retain strong hydrogen lines and arise from stars that preserve their outer hydrogen envelopes.

Type Ib and Ic supernovae lack hydrogen and, in the latter case, helium signatures, indicating extensive pre-explosion mass loss. These core-collapse events produce irregular light curves and depend sensitively on progenitor properties.

Type Ia supernovae originate from a different mechanism. These events occur when a white dwarf in a binary system accretes matter from a companion and approaches the Chandrasekhar mass. Carbon and oxygen ignite under degenerate conditions, completely disrupting the white dwarf. Unlike core-collapse events, Type Ia supernovae leave no remnant. Their uniformity leads to a characteristic light curve with a distinctive rise and decay pattern, allowing for empirical standardisation and making them cosmological distance indicators.

Type Ia supernovae serve as standard candles with intrinsic luminosity L inferred after empirical correction. Observers measure apparent flux F , related by $F = L/(4\pi d_L^2)$ where d_L is luminosity distance. Astronomers use magnitudes: apparent magnitude m minus absolute magnitude M (at 10 parsecs, where one parsec is approximately 3.26 light-years and 1 Mpc = 10^6 parsecs) gives the distance modulus $\mu = m - M = 5 \log_{10}(d_L/\text{Mpc}) + 25$. After correcting for light curve shape (brighter events fade slower) and extinction, standardised magnitudes yield cosmological information.

Each supernova provides two observables: a redshift z and a distance modulus $\mu(z)$. The redshift is defined by $1 + z = \frac{\lambda_{\text{obs}}}{\lambda_{\text{emit}}}$ where λ_{emit} is the rest-frame emission wavelength and λ_{obs} is the observed wavelength. At cosmological scales this reflects the changing scale factor. The set of points (z, μ) defines a Hubble diagram, which records the history of cosmic expansion.

In the 1990s, two major collaborations used wide-field imaging and spectroscopy to discover Type Ia supernovae out to $z \sim 1$, coordinating ground-based and Hubble telescopes. Each event was monitored across bands, with light curves compared to templates and corrected for extinction and host properties.

Distant supernovae were fainter than expected (Riess et al., 1998) for a decelerating universe. The effect persisted across all filters, instruments, and analysis methods. Cross-calibration against low-redshift samples and repeated observations supported the result. The implication was that the universe had expanded more during the photons' travel than gravitational deceleration would allow, requiring an expansion rate that increases with time.

Cosmic evolution follows the scale factor $a(t)$ —expansion occurs when $a(t)$ increases, acceleration when $\ddot{a}(t) > 0$. Matter and radiation cause deceleration (positive density, non-negative pressure), so acceleration requires a component with sufficiently negative pressure. Supernovae revealed this directly: $d_L(z)$ increases faster with redshift than expected for deceleration, following purely from empirical brightness-distance relations.

The cosmic microwave background (CMB) confirmed this: temperature fluctuations indicate flat geometry, but matter provides only one-third of the required density. The remainder must be a non-clustering, smooth component. Large-scale clustering observations, including baryon acoustic oscillations, support this picture (Eisenstein et al., 2005).

Type Ia supernovae first detected acceleration; CMB and structure surveys confirmed it. The converging evidence strongly favours an accelerating expansion history.

The possibility of a component that modifies expansion on large scales had been considered decades earlier. Einstein added Λ to his field equations in 1917 to construct a static universe, with repulsion balancing gravity. This static solution was unstable and soon discarded after Hubble's 1929 observations demonstrated that galaxies are receding from one another. The cosmological constant was removed from most models and became a historical footnote.

Its mathematical form, however, remained valid. Λ represents a constant energy density that does not dilute as the universe expands. In general relativity, it contributes with a sign and magnitude that counteract the deceleration caused by matter. In quantum field theory, the vacuum itself carries an energy density that enters the gravitational equations in the same way. Estimates of this vacuum energy, however, exceed observed values by many orders of magnitude, creating a mismatch that remains unresolved.

When supernovae revealed acceleration, the cosmological constant was reintroduced. Its inclusion adjusts the predicted luminosity-distance curve to align with observations.

'Dark energy' drives cosmic acceleration with behaviour distinct from matter or radiation—it doesn't cluster and maintains constant energy density as space expands. In its simplest form, it corresponds to Λ , though models allow time variation via scalar fields or modified gravity. Observationally, it's defined by producing acceleration: models excluding it fail to match supernovae, CMB patterns, and clustering measurements.

The inference of cosmic acceleration relies on systematic measurement of distance and redshift. Each supernova provides a data point (z, μ) on the Hubble diagram, building one of the most important experimental datasets for cosmological models.

Surveys use low-redshift controls and Cepheid calibration (standard candles) to eliminate systematics. Large surveys (SNLS, SDSS-II, DES) compiled thousands of supernovae to $z \sim 1.5$, confirming acceleration—the Hubble diagram curvature cannot be replicated by matter-only models. Independent measurements from baryon acoustic oscillations serve as standard rulers across redshift, further constraining the expansion history.

Combined data constrain acceleration's evolution. Future surveys—including the Vera Rubin LSST (first data 2025) and the Nancy Grace Roman Space Telescope (expected 2027)—aim to refine measurements of $d_L(z)$ to sub-percent precision, distinguishing between a true cosmological constant and time-varying models of dark energy.

The cosmological constant Λ provides the simplest model for dark energy. Within general relativity, cosmic expansion follows the Friedmann equations (Friedmann, 1922) governing the scale factor $a(t)$ —comoving separations grow as $a(t)$ increases. It can be shown that acceleration depends on energy density and pressure: ordinary matter and radiation (non-negative pressure) cause deceleration, while components with sufficiently negative pressure drive acceleration. The cosmological constant corresponds to uniform energy density with negative pressure equal to its energy density, acting as a repulsive gravitational source that accelerates expansion.

Combining Λ with cold dark matter—non-relativistic, weakly interacting matter that clusters gravitationally—produces Λ CDM, the standard cosmological model. It successfully describes a wide range of observations, placing the universe's current energy budget at

68% dark energy, 27% dark matter, and 5% ordinary matter—consistent with supernovae, CMB, and structure observations.

High-redshift supernovae confirm a transition from an earlier decelerating phase, when matter dominated the energy budget, to the current accelerating phase driven by dark energy. This transition occurred roughly when the universe was half its current age, when the energy densities of matter and dark energy became comparable.

The physical nature of dark energy remains unknown. The observed value of the cosmological constant is extraordinarily small compared to theoretical expectations from quantum field theory. Quantum mechanics predicts that even empty space should possess a vacuum energy density, arising from zero-point fluctuations of all possible fields. Yet this predicted vacuum energy exceeds the observed dark energy density by approximately 120 orders of magnitude—a discrepancy so severe it has been called ‘the worst theoretical prediction in the history of physics’ (Hobson, Efstathiou and Lasenby, 2006, p. xi; see also the chapter epigraph)

This cosmological constant problem exposes an incomplete understanding of quantum gravity. General relativity couples to absolute energy density, not just differences. The vacuum energy cannot simply be subtracted away without affecting the geometry of space and time. The resolution may require entirely new physics beyond the Standard Model or a radical revision of our understanding of gravity.

How Much Can Be Measured

Scientific understanding is constrained by what can be observed. In cosmology, measurements rely on electromagnetic radiation, primarily in a few accessible bands of the spectrum. Telescopes capture only a fraction of the sky at any given time. Most stars, galaxies, and intergalactic matter are never directly observed, and yet, from this limited sample, consistent physical laws have been extracted.

This is a serendipitous outcome! Sampling 0.0...01% of the universe's contents has been sufficient to uncover mathematical relationships that describe its large-scale behaviour. The same formalism that governs an apple's descent also predicts planetary motion, galaxy trajectories, and the expansion of space.

The precision is remarkable: measuring a falling spoon on Earth can predict planetary orbits in distant galaxies with 99% accuracy. Adding the considerable effort of developing and testing general relativity gains the remaining 0.99% precision needed for GPS satellites and gravitational wave detectors. Yet despite this extraordinary success in describing gravity and motion across cosmic scales, 95% of the total energy content in the universe remains completely unaccounted for!

These components dominate the dynamics. The success of mathematical modelling in organising what is accessible illustrates both the remarkable reach of inference from incomplete information and the humbling limits of our cosmic perspective.

The Custodians of One

The monastery on Tethra-9 had no name, it needed none. Carved from igneous glass in a moon's crust whose atmosphere had boiled off in the Proxima Flare, it housed monks who had transcended insignia, watching null-entropy with infinite patience.

Commander Rafe Lin arrived in *Shibboleth*, a vacuum-energy cruiser extracting work from the ground state itself. Its hull was reinforced against gradients that could collapse neutron stars into geometric points—an abomination, even to its builders.

The monks observed his approach with studied indifference. Seven levels down through recursive dimensional barriers: a featureless black sphere hovering in defiance of every conservation law except the one that mattered.

The Monad. A magnetic monopole—not merely a particle, but the *arithmetic* upon which particle physics balanced its ledgers. Without it, electric charge would be as arbitrary as poetry, gauge symmetries unravelling like a bad proof.

“Extraction protocol, immediate,” Rafe commanded, his voice carrying the authority of those who were never disobeyed.

The eldest monk spoke with the gentleness reserved for those about to discover the difference between locally and globally defined: “Commander, your vessel maintains structural integrity only because spacetime agrees to be continuous. Remove the boundary condition, and agreement becomes... negotiable.”

The containment field collapsed. The sphere shifted through the space of *definitions* themselves—somewhere, the Riemann hypothesis lost a prime, fundamental constants performed a probabilistic quadrille.

Shibboleth discovered its existence had been predicated on several no-longer-valid assumptions. The ship lost coherence with the dignity of a well-posed problem becoming ill-defined. Rafe learned, too late, that identity is a function of existence.

The monks reactivated containment with the practised efficiency of librarians re-shelving infinity. The sphere returned to its fixed point. The universe exhaled. No rescue missions were authorised, for there was nothing left to rescue but mathematics.

And the cosmos, whole once more, continued its stately rotation, because $qg = \frac{n\hbar}{2}$ only holds if $\oint_{\partial V} \mathbf{B} \cdot d\mathbf{A} \neq 0$ somewhere.

The Cosmological Constant Problem: Quantum Vacuum Energy vs Observations

Quantum Field Theory Prediction

In quantum field theory, even empty space possesses energy due to zero-point fluctuations. For a free massless scalar field $\phi(x, t)$ with Hamiltonian,

$$H = (1/2) \int [\pi^2(x) + |\nabla\phi(x)|^2] d^3x,$$

Upon canonical quantization, the vacuum expectation value becomes:

$$\langle 0|H|0\rangle = (1/2) \int d^3k/(2\pi)^3 \cdot \hbar\omega_k,$$

where $\omega_k = c|\mathbf{k}|$. This yields:

$$\langle 0|H|0\rangle \propto \int_0^\infty k^3 dk,$$

which diverges as k^4 at high momentum, requiring a cutoff.

Planck-Scale Cutoff

Assuming quantum field theory remains valid up to the Planck energy scale:

$$k_{\max} = M_{\text{Planck}}c/\hbar = \sqrt{c^3/(\hbar G)}.$$

The vacuum energy density becomes:

$$\begin{aligned} \rho_{\text{vac}}^{\text{theory}} &= \hbar c/(16\pi^2) \cdot k_{\max}^4 \\ &= \hbar c/(16\pi^2) \cdot (c^6/(\hbar^2 G^2)) \\ &= c^7/(16\pi^2 \hbar G^2) \sim 10^{76} \text{ GeV}^4. \end{aligned}$$

Observational Constraints

Cosmological observations from supernovae, CMB, and large-scale structure constrain the dark energy density to:

$$\rho_{\text{DE}}^{\text{obs}} = \rho_{\text{crit}}\Omega_\Lambda \approx (3H_0^2/(8\pi G)) \times 0.68,$$

where $H_0 \approx 70 \text{ km/s/Mpc}$. Converting to natural units:

$$\rho_{\text{DE}}^{\text{obs}} \approx 10^{-47} \text{ GeV}^4.$$

The Discrepancy

The ratio of theoretical prediction to observational constraint:

$$\rho_{\text{vac}}^{\text{theory}}/\rho_{\text{DE}}^{\text{obs}} \sim 10^{76}/10^{-47} = 10^{123}.$$

This represents the largest mismatch between theory and observation in physics history. (Depending on cutoff choices and degrees of freedom, values from 10^{118} to 10^{123} are quoted in the literature.)

The Fine-Tuning Problem

Unlike other physics areas where only energy differences matter, general relativity couples directly to absolute energy density through Einstein's field equations:

$$G_{\mu\nu} + \Lambda g_{\mu\nu} = (8\pi G/c^4)T_{\mu\nu},$$

where $\Lambda = (8\pi G/c^2)\rho_{\text{vac}}$.

If quantum vacuum energy contributed at the predicted level, it would drive exponential expansion so rapid that structure formation would be impossible. The observed value requires either:

1. Extraordinary cancellation reducing vacuum energy by 120 orders of magnitude.
2. New physics beyond the Standard Model altering vacuum structure.
3. Modification of general relativity at cosmological scales.

No proposed solution has gained broad acceptance, making this one of the most pressing problems in theoretical physics.

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